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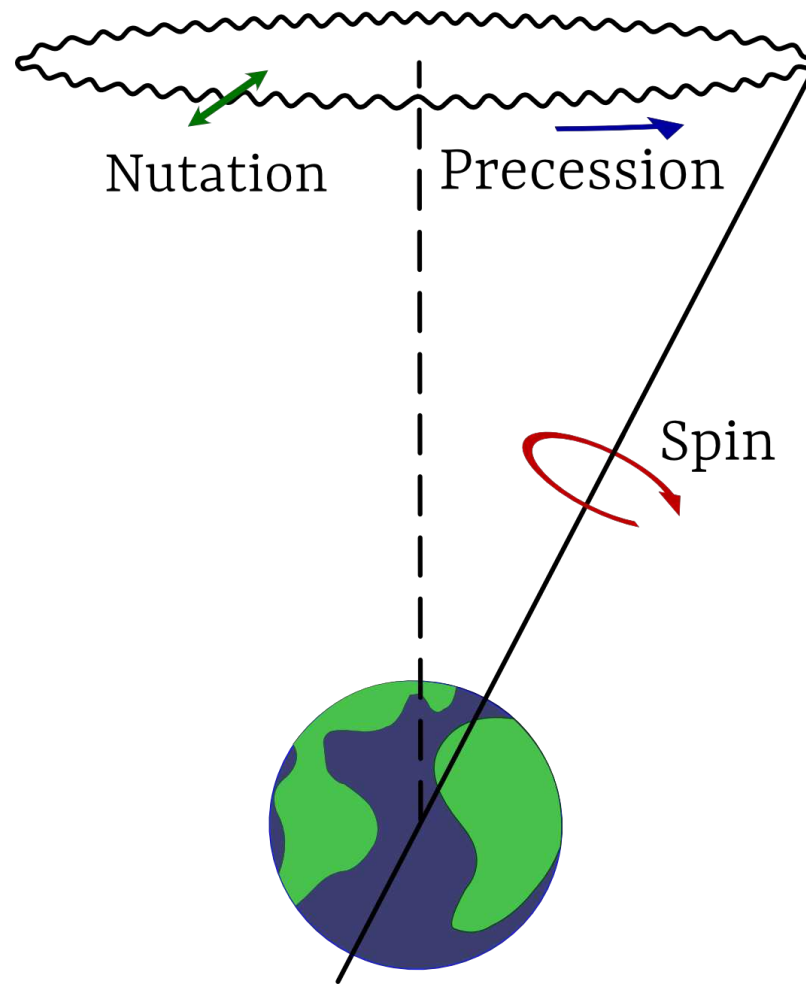
Topographic effects in a stratified layer at the top of the core

IAGA-IASPEI Virtual Conference, 21st – 27th August 2021

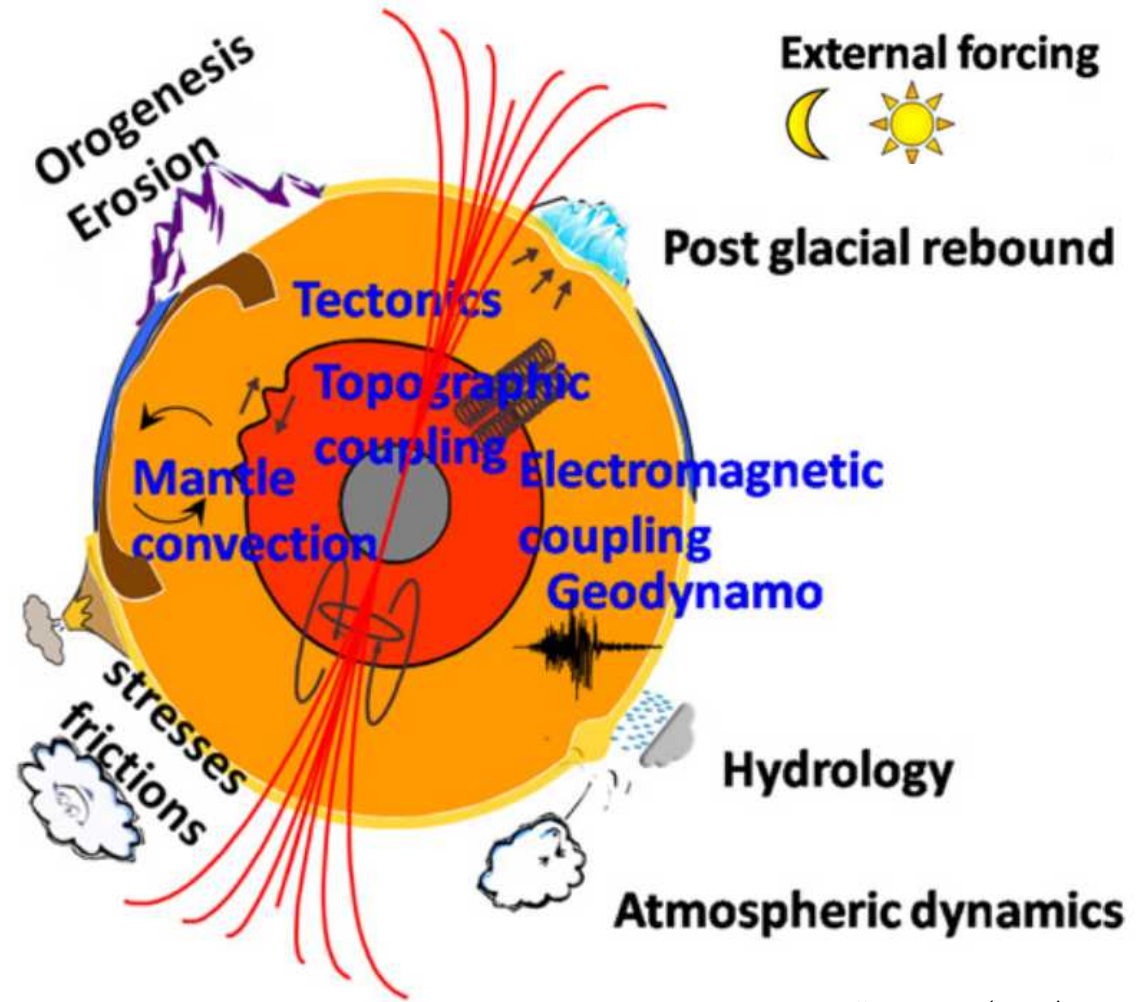
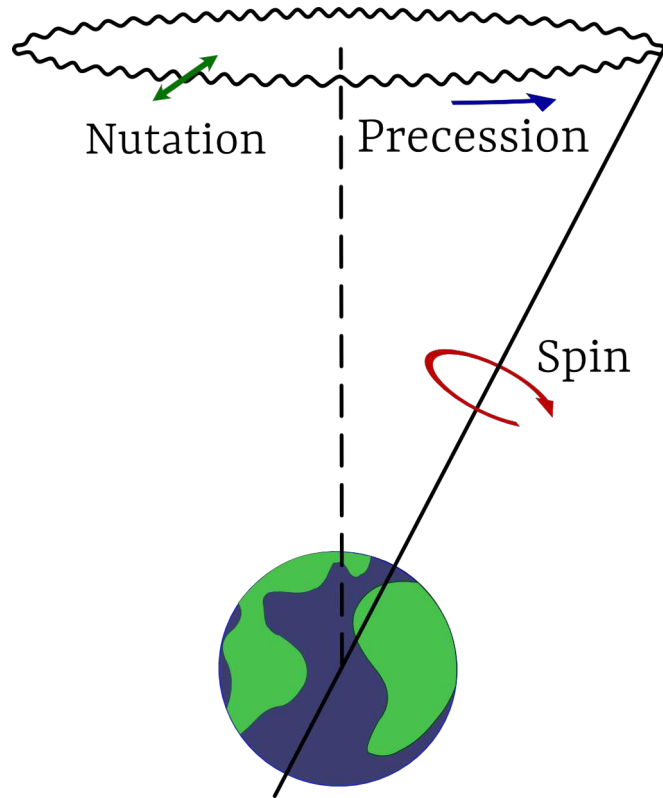
Rémy Monville, David Cébron, Dominique Jault – 2021

ISTerre, Université Grenoble Alpes, CNRS

Variations of the Earth's rotation



Variations of the Earth's rotation



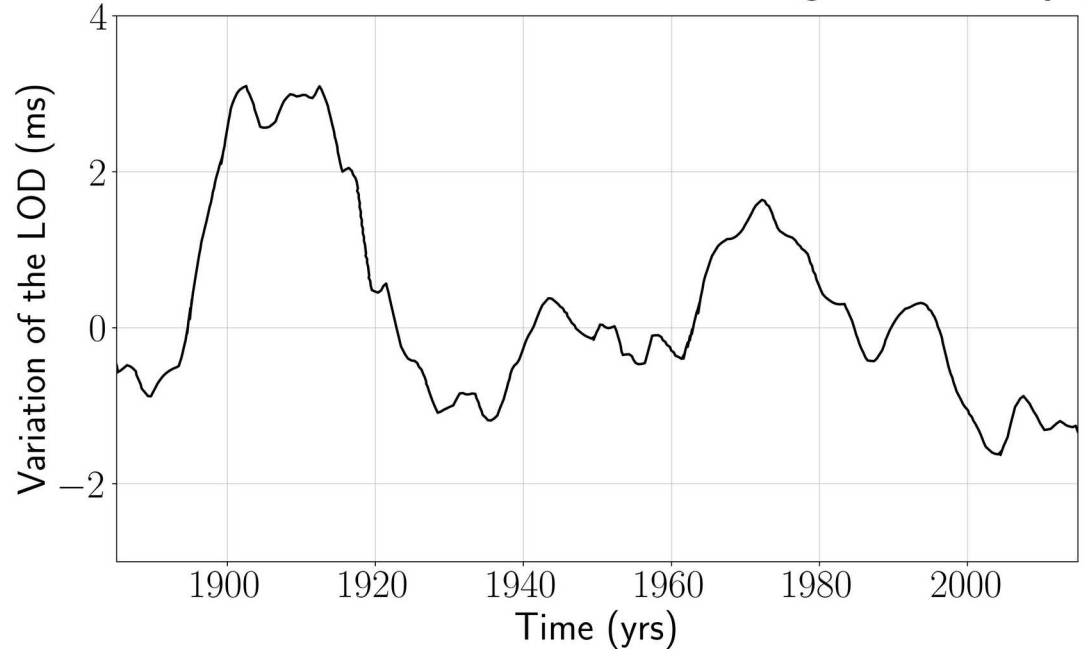
Karatekin, Ö. et al. (2011)

Exchanges of angular momentum

Core coupling:

- Electromagnetic
- Viscous
- Gravitational
- Pressure torque on small scale topography

Core contribution to the variation of the length of the day (LOD)



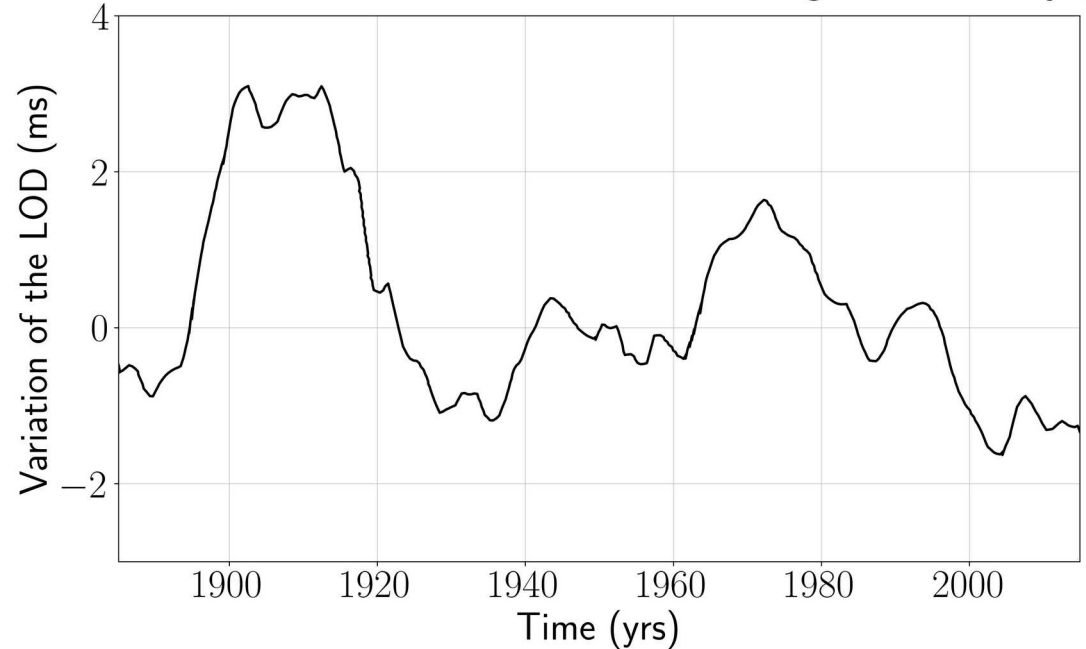
Gillet, et al. (2019).

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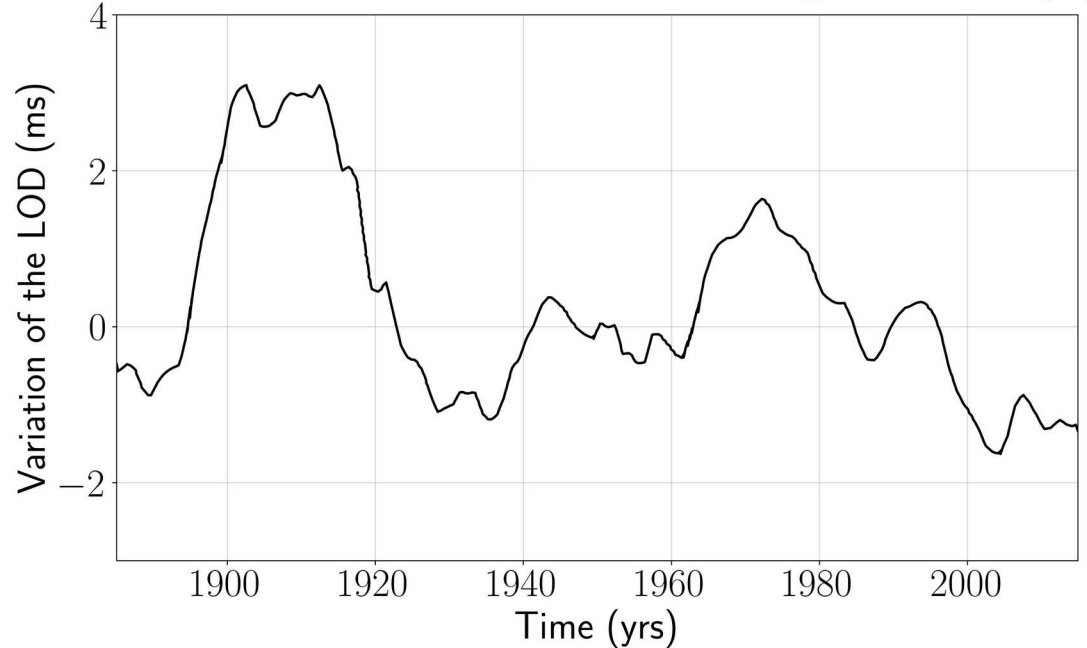
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Exchanges of angular momentum

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Gillet, et al. (2019).

Motivations

Can the **small scale** topographic coupling explain:

- The decadal changes in the **Length-of-Day** (*Glane and Buffett 2018, Jault 2020*) ?
- The out of phase component of the retrograde annual **nutations** of the Earth's rotation axis (*Buffett 2010*) ?

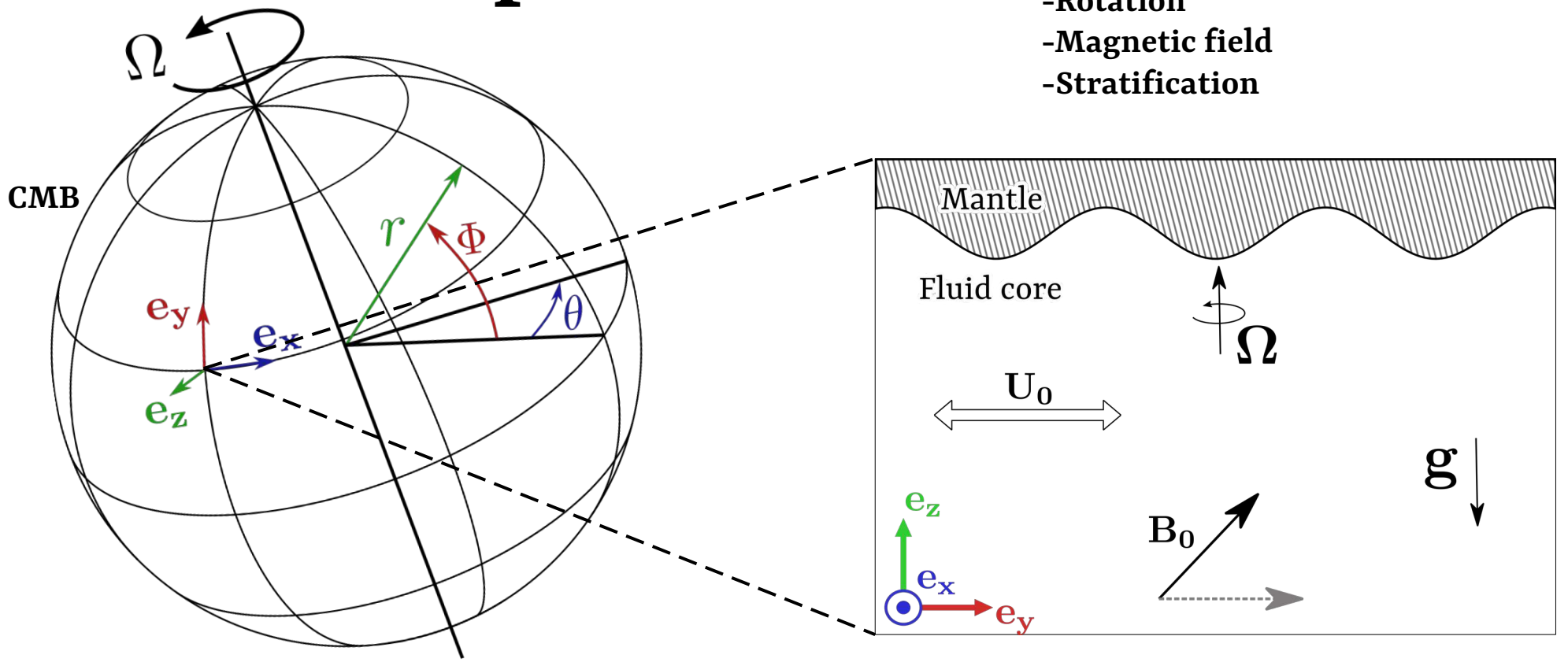
How well can a **local perturbative model** help us to understand these measurements, and what are its limitations ?

Geometry of the problem

Spherical shell \rightarrow Cartesian frame

Key effects:

- Rotation
- Magnetic field
- Stratification



Geometry of the problem

Basic State:

- **Velocity:**

$$\mathbf{U}_0 = u_0 \Re_e (e^{I\omega t}) \mathbf{e}_x$$

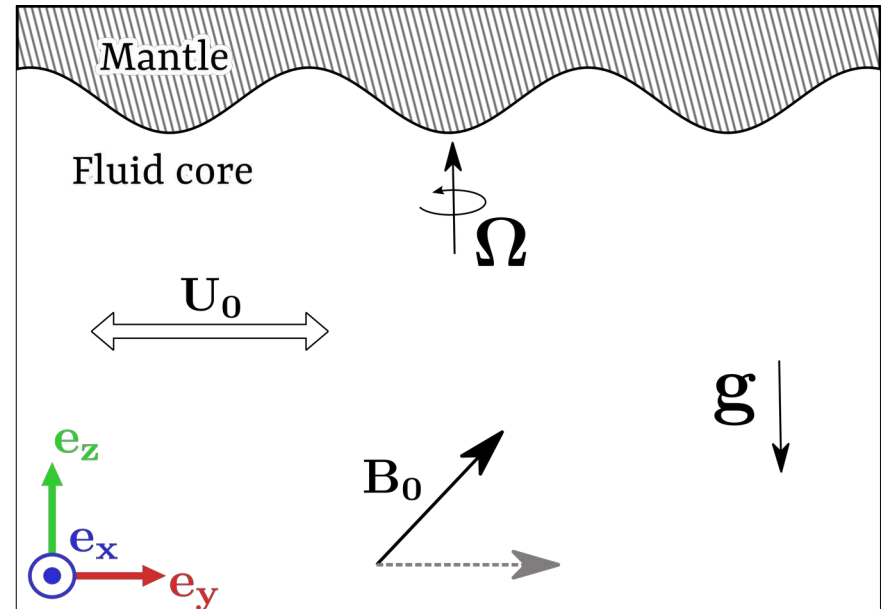
or \rightarrow

- **Magnetic field:**

$$\mathbf{B}_0 = b_0 (\cos(\Phi) \mathbf{e}_y - 2 \sin(\Phi) \mathbf{e}_z)$$

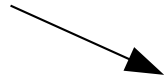
- **Density:**

$$\rho = \rho_r (1 - \alpha z)$$



Equations of motion

Navier-Stokes



$$\rho_r D_t \mathbf{U} = -\rho_r \chi \mathbf{e}_z \times \mathbf{U} - \nabla p + \mu \nabla^2 \mathbf{U} + \rho \mathbf{g} + \mathbf{J} \times \mathbf{B},$$

Mass conservation $\longrightarrow \partial_t \rho + (\mathbf{U} \cdot \nabla) \rho = 0,$

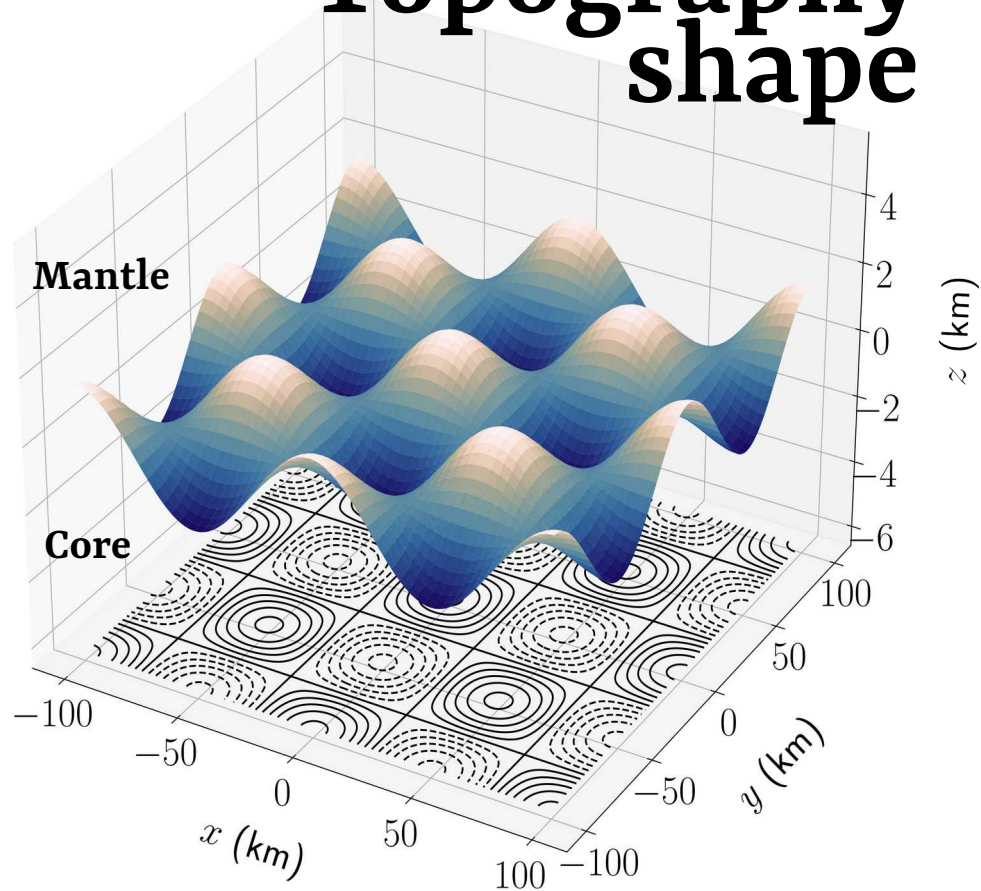
Magnetic field
(Induction) $\longrightarrow \partial_t \mathbf{B} = \eta \nabla^2 \mathbf{B} + \nabla \times (\mathbf{U} \times \mathbf{B}),$

Incompressible fluid $\longrightarrow \nabla \cdot \mathbf{U} = 0,$

Conservation of magnetic flux $\longrightarrow \nabla \cdot \mathbf{B} = 0,$

Magneto-hydro-dynamic equations (MHD), in Boussinesq approximation

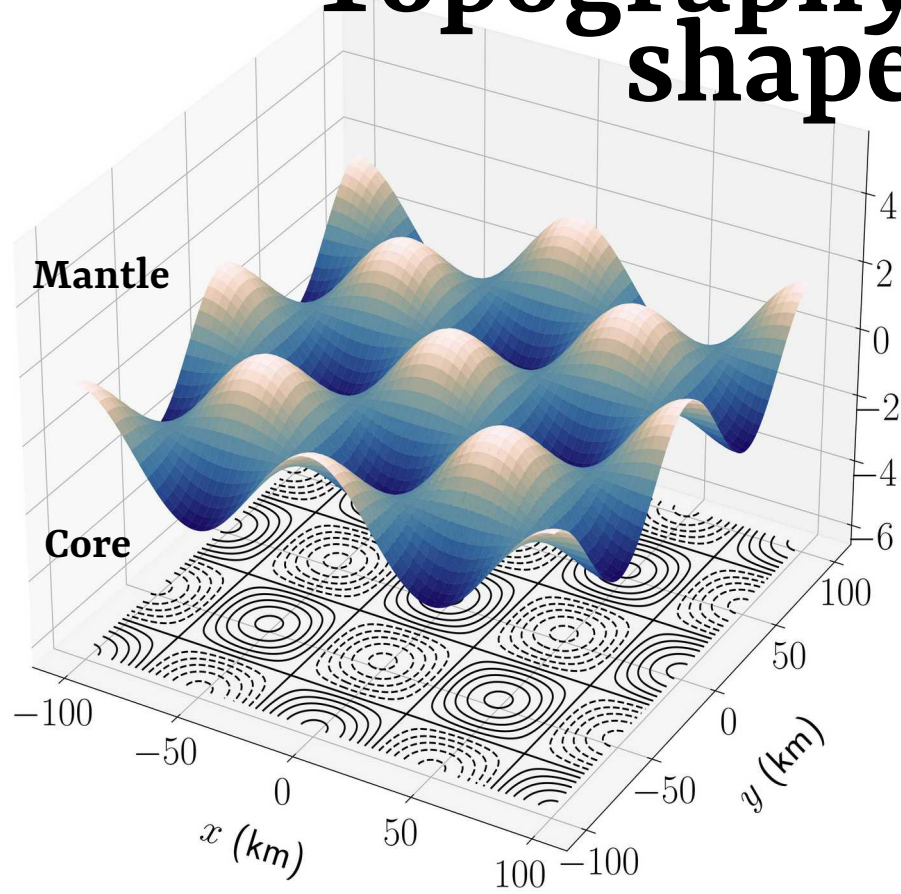
Topography shape



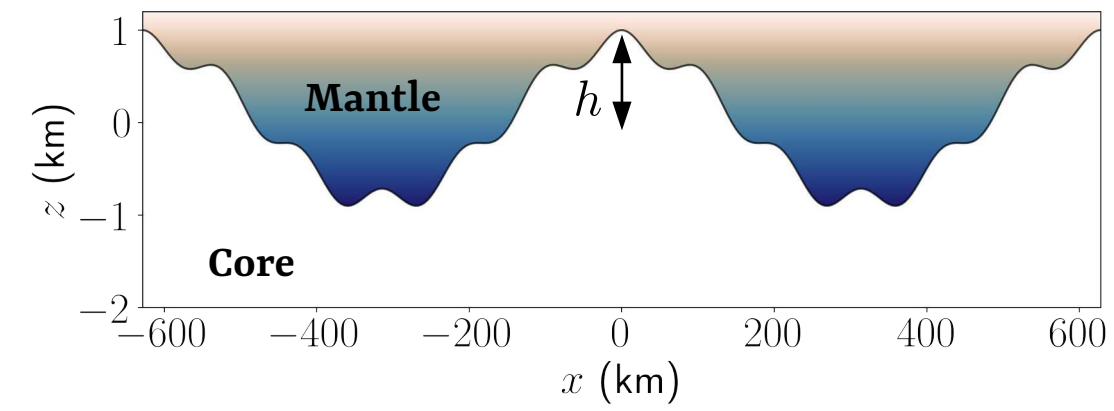
$$h(x, y) = \epsilon_t \sum_{j=0}^n \Re(\exp(\mathbf{i} \mathbf{k}_j \cdot \mathbf{r})) / n,$$

$\epsilon_t \ll 1$ is the topography height divided by a typical length scale

Topography shape

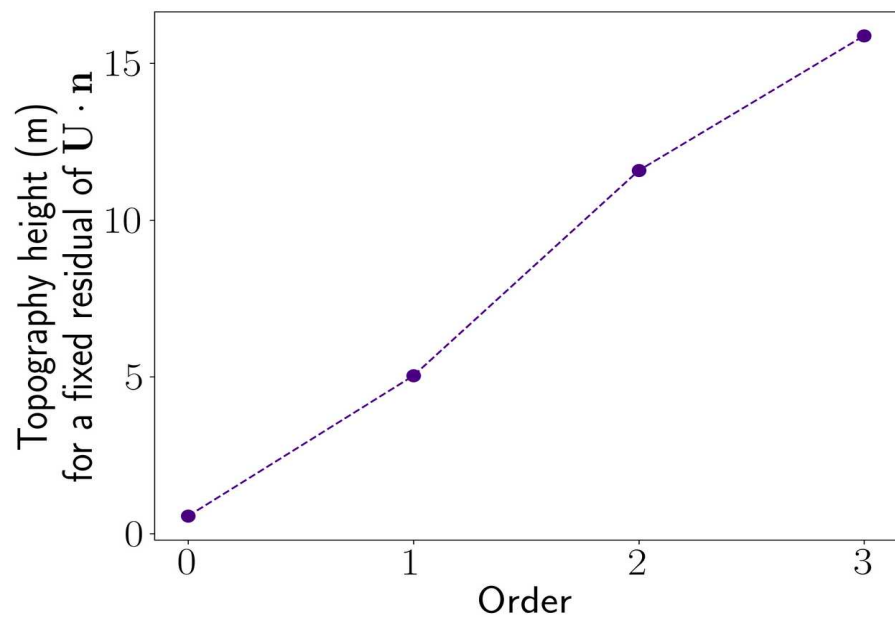


$$h(x, y) = \epsilon_t \sum_{j=0}^n \Re(\exp(i\mathbf{k}_j \cdot \mathbf{r})) / n,$$



Methods

Solving equations with with a weakly non linear perturbation approach

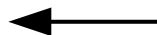


Limited by a small parameter

→ Topography height

Glane and Buffett (2018) : ~30-50m

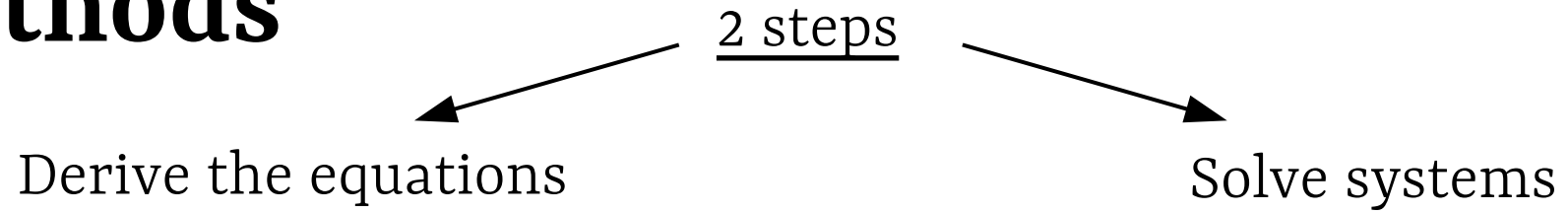
Buffett (2010) : ~100m



Higher orders of perturbation

Quasi linear variation when the serie is convergent

Methods

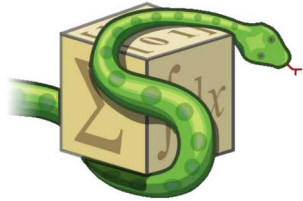


Methods

2 steps

Derive the equations

Solve systems



Sympy : *symbolic mathematics*



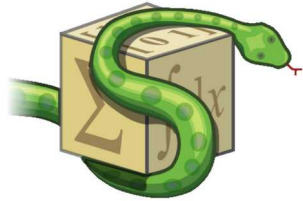
mpmath : *arithmetic with arbitrary precision*

Methods

2 steps

Derive the equations

Solve systems



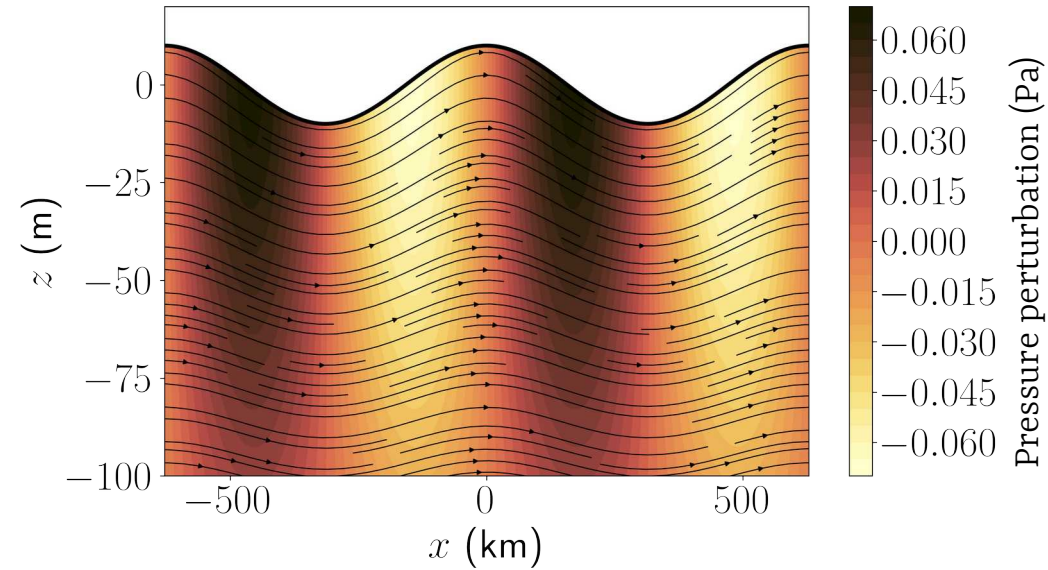
Sympy : symbolic mathematics

python™

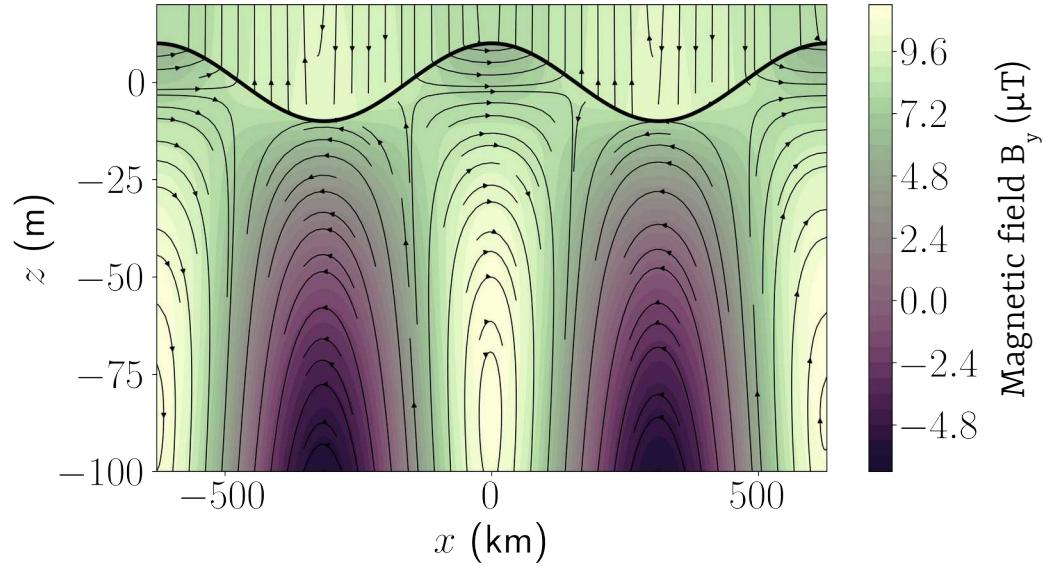
mpmath : arithmetic with arbitrary precision

My code : **ToCCo**

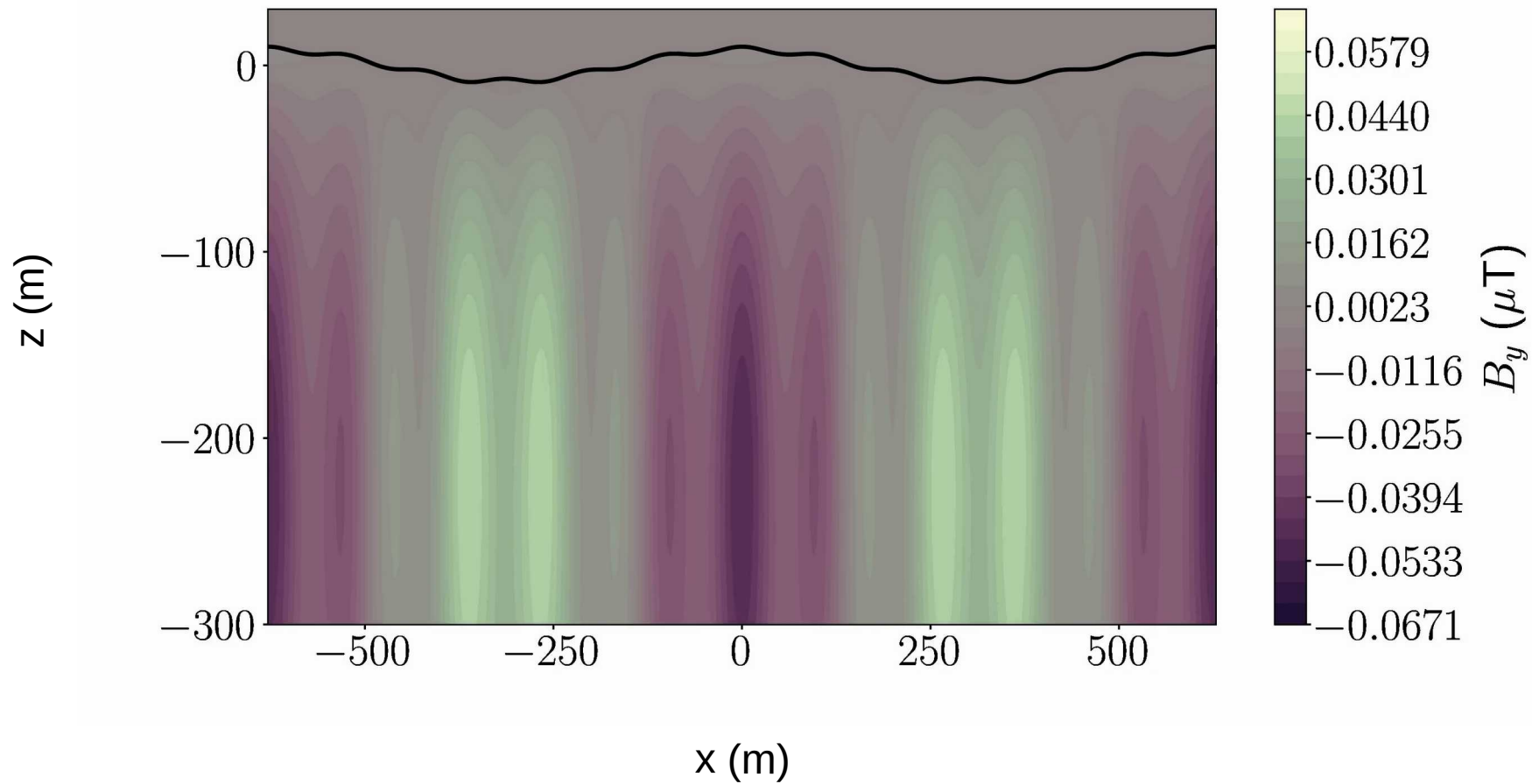
Results



Streamlines = velocity



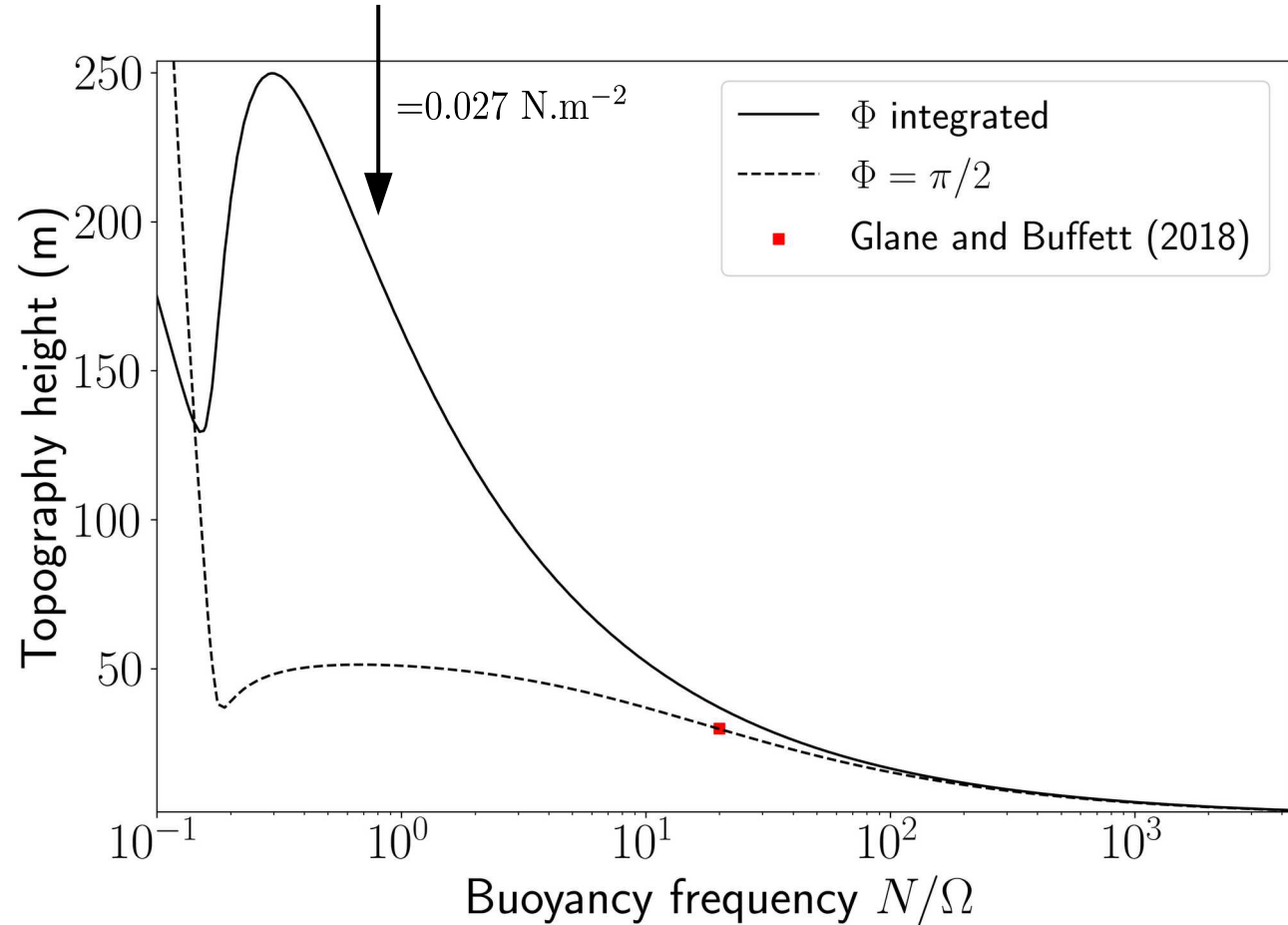
Streamlines = Current density



What is the value for (h, N) required to explain the observed variation of the length of the day ?

- Steady and uniform flow
- Insulating mantle

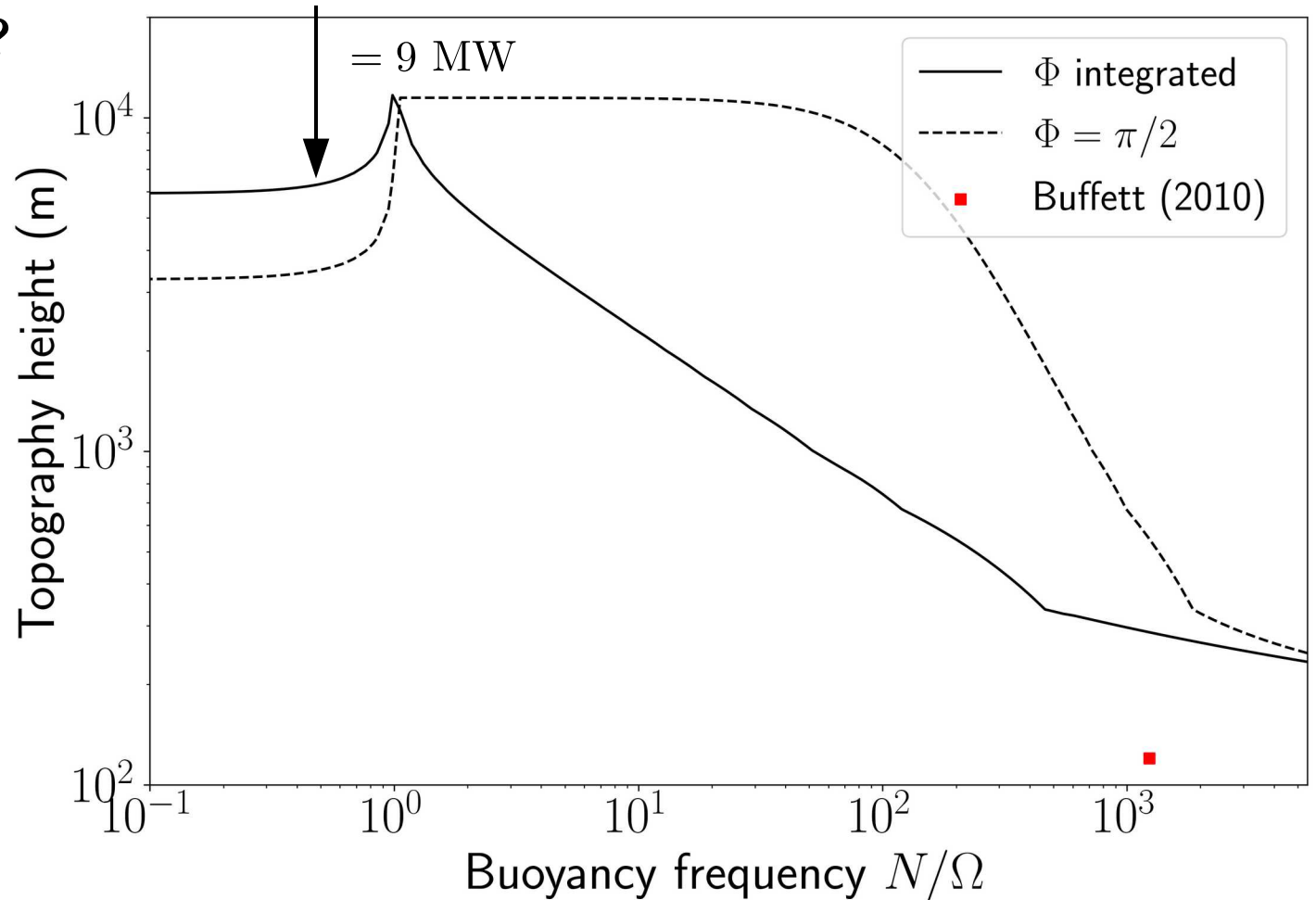
integrated = integration with latitude, taking into account the variation of Ω and \mathbf{B}_0



What is the value for (h,N) required to explain the observed dissipative coupling ?

- Oscillating flow with diurnal period
- Conducting mantle
→ electrical conductivity ratio : $\frac{\sigma_{core}}{\sigma_{mantle}} = 500$
- At the pole : $B_0 = 0.5 \text{ mT}$

integrated = integration with latitude, taking into account the variation of Ω and \mathbf{B}_0

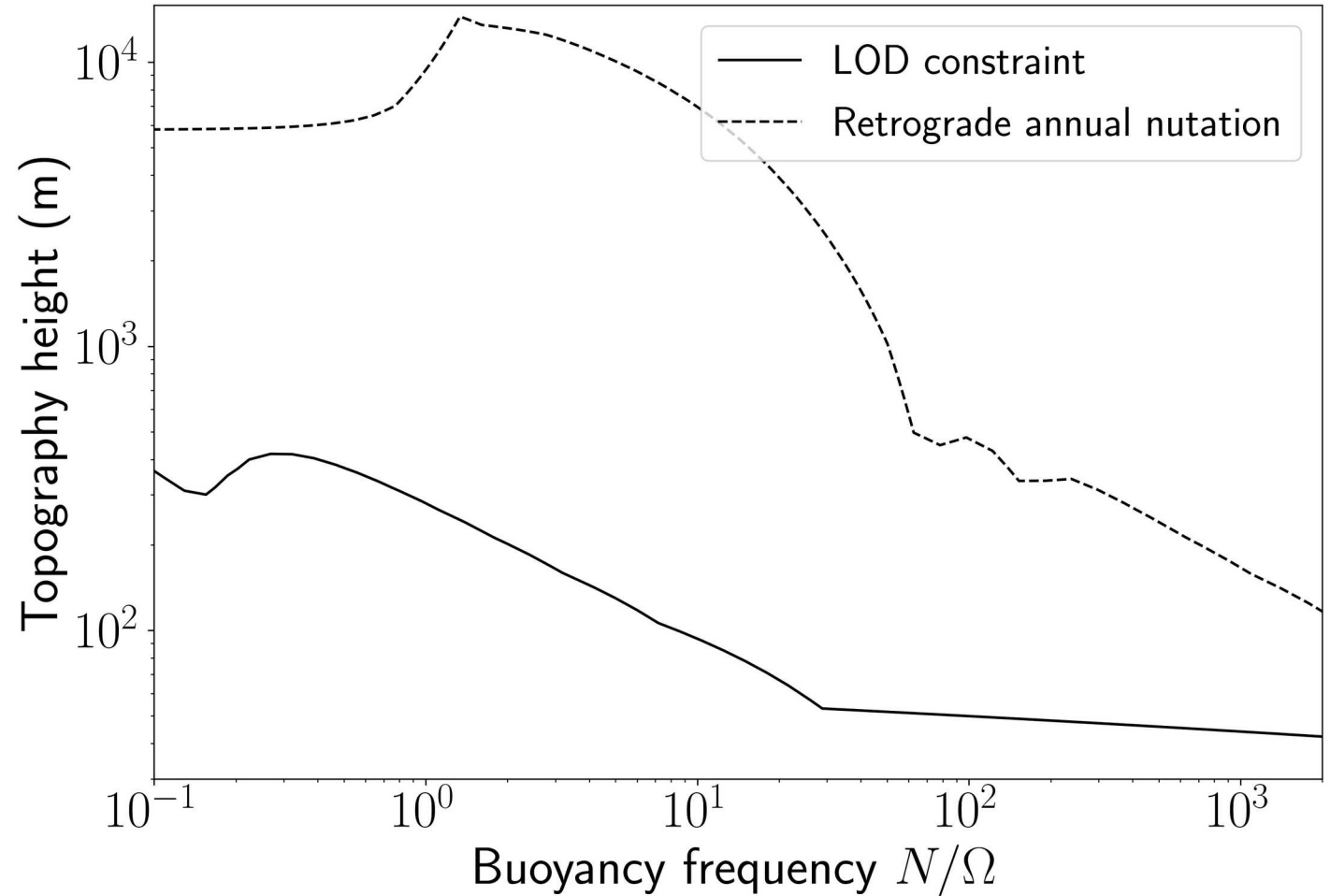


Comparing at the same parameters

-Integrated with latitude

-Conducting mantle
→ electrical conductivity
ratio : $\frac{\sigma_{core}}{\sigma_{mantle}} = 500$

- At the pole : $B_0 = 0.65$ mT



Conclusion

- We can, with our model, explore a wide panel of parameters in a consistent manner
- developed at a higher order of perturbation
- Our model, in its simplest form, does not explain simultaneously the two sets of data (LOD, nutation)

and Perspectives

- Couple our results with Earth model of rotation
- Study the convergence radius of the perturbative model and constrain its limit of applicability.

References

B. A. Buffett, “Chemical stratification at the top of Earthcore: Constraints from observations of nutations,” *Earth and Planetary Science Letters*, vol. 296, no. 3-4, Art. no. 3-4, Aug. 2010, doi: 10.1016/j.epsl.2010.05.020.

N. Gillet, L. Huder, and J. Aubert, “A reduced stochastic model of core surface dynamics based on geodynamo simulations,” *Geophysical Journal International*, vol. 219, no. 1, Art. no. 1, 2019.

S. Glane and B. Buffett, “Enhanced Core-Mantle Coupling Due to Stratification at the Top of the Core,” *Frontiers in Earth Science*, vol. 6, 2018, doi: 10.3389/feart.2018.00171.

D. Jault, “Tangential stress at the core–mantle interface,” *Geophysical Journal International*, vol. 221, no. 2, Art. no. 2, Jan. 2020, doi: 10.1093/gji/ggaa048.

Ö. Karatekin et al., “Atmospheric angular momentum variations of Earth, Mars and Venus at seasonal time scales,” *Planetary and Space Science*, vol. 59, no. 10, Art. no. 10, 2011.

Equations of motion

Inviscid fluid

Inertia
 Coriolis force (with Beta-plane approximation)
 Reduced pressure gradient
 Buoyancy
 Lorentz force

$$\boxed{\rho_r D_t \mathbf{U}} = - \boxed{\rho_r \chi \mathbf{e}_z \times \mathbf{U}} - \boxed{\nabla p} + \cancel{\mu \nabla^2 \mathbf{U}} + \boxed{\rho \mathbf{g}} + \boxed{\mathbf{J} \times \mathbf{B}},$$

$$\partial_t \rho + (\mathbf{U} \cdot \nabla) \rho = 0,$$

$$\partial_t \mathbf{B} = \boxed{\eta \nabla^2 \mathbf{B}} + \boxed{\nabla \times (\mathbf{U} \times \mathbf{B})},$$

Joule Dissipation
 Coupling term
 $\nabla \cdot \mathbf{U} = 0,$

Magneto-hydro-dynamic equations (MHD)